Asymptotic stability of kinetic plasmas for general collision potentials

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Background

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Physical description of a plasma

- ► Plasma is the 4th state of matter: solid→liquid→gas→plasma
- ▶ 99.9% of the universe exists in a plasma state
- Plasma is a gas of charged particles, e.g. electrons and ions
- The motion of plasmas strongly responds to the self-consistent electromagnetic field through the Maxwell equations

$$\frac{1}{c}\partial_t E - \nabla \times B = -\frac{4\pi}{c}J, \quad \frac{1}{c}\partial_t B + \nabla \times E = 0,$$
$$\nabla \cdot E = 4\pi\rho, \quad \nabla \cdot B = 0.$$

- Plasma physics involves the physics of classical mechanics, electromagnetism, and non relativistic statistical mechanics
- ► Challenge lies in the long-range coulomb interaction

Mathematical description of a plasma

- microscopic particle model for $[x_i(t), \xi_i(t)]$
- mesoscopic kinetic model for $f(t, x, \xi)$
- macroscopic fluid model for [n(t, x), u(t, x)]

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1st type (Klimontovich)

Microscopic motion equations governing $[x_i(t), \xi_i(t)]$ of all plasma particles $1 \le i \le N_0$ of *s*-species at any time *t*:

$$\begin{split} m_s \frac{d\xi_i}{dt} &= q_s [E(t, x_i) + \frac{\xi_i}{c} \times B(t, x_i)], \\ \frac{1}{c} \partial_t E - \nabla \times B &= -\frac{4\pi}{c} J, \\ \frac{1}{c} \partial_t B + \nabla \times E &= 0, \\ \nabla \cdot E &= 4\pi\rho, \quad \nabla \cdot B &= 0, \\ \rho &= \sum_s q_s \int_{\mathbb{R}^3} N_s(t, x, \xi) \, d\xi, \quad J &= \sum_s q_s \int_{\mathbb{R}^3} \xi N_s(t, x, \xi) \, d\xi, \\ N_s(t, x, \xi) &= \sum_{i=1}^{N_0} \delta(x - x_i(t)) \delta(\xi - \xi_i(t)). \end{split}$$

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2nd type (kinetic plasma equations)

$$\begin{split} f_s &= f_s(t, x, \xi), \ t \ge 0, x \in \mathbb{R}^3, \xi \in \mathbb{R}^3, \ s = i, e \\ &\frac{\partial_t f_s + \xi \cdot \nabla_x f_s + \frac{q_s}{m_s} (E + \frac{\xi}{c} \times B) \cdot \nabla_\xi f_s = \left(\frac{\partial f_s}{\partial t}\right)_c, \\ &\frac{1}{c} \partial_t E - \nabla \times B = -\frac{4\pi}{c} J, \\ &\frac{1}{c} \partial_t B + \nabla \times E = 0, \\ &\nabla \cdot E = 4\pi\rho, \quad \nabla \cdot B = 0, \\ &\rho = \sum_s q_s \int_{\mathbb{R}^3} f_s(t, x, \xi) \, d\xi, \quad J = \sum_s q_s \int_{\mathbb{R}^3} \xi f_s(t, x, \xi) \, d\xi. \end{split}$$

Depending on the collisional feature, the system is called

- Vlasov-Maxwell-Boltzmann
- Vlasov-Maxwell-Landau

Characterization of collisions

- Boltzmann collision (Boltzmann, 1872)

$$\begin{pmatrix} \frac{\partial f_s}{\partial t} \end{pmatrix}_c = \sum_{s'} Q(f_s, f_{s'}),$$

$$Q(f_1, f_2)(\boldsymbol{\xi}) = \iint_{\mathbb{R}^3 \times S^2} B(\boldsymbol{\xi} - \boldsymbol{\xi}_*, \omega) \{f_1(\boldsymbol{\xi}') f_2(\boldsymbol{\xi}'_*) - f_1(\boldsymbol{\xi}) f_2(\boldsymbol{\xi}_*)\} d\boldsymbol{\xi}_* d\omega,$$

$$\begin{cases} \boldsymbol{\xi}' = \boldsymbol{\xi} & -\frac{2m_2}{m_1 + m_2} [(\boldsymbol{\xi} - \boldsymbol{\xi}_*) \cdot \omega] \omega, \\ \boldsymbol{\xi}'_* = \boldsymbol{\xi}_* + \frac{2m_1}{m_1 + m_2} [(\boldsymbol{\xi} - \boldsymbol{\xi}_*) \cdot \omega] \omega, \\ B(\boldsymbol{\xi} - \boldsymbol{\xi}_*, \omega) = \Phi(|\boldsymbol{\xi} - \boldsymbol{\xi}_*|) b(\frac{\boldsymbol{\xi} - \boldsymbol{\xi}_*}{|\boldsymbol{\xi} - \boldsymbol{\xi}_*|} \cdot \omega),$$

$$\Phi(|\boldsymbol{z}|) \sim |\boldsymbol{z}|^{\gamma} (-3 < \gamma \le 1), \quad \sin \theta b(\cos \theta) \sim \frac{1}{\theta^{1+2\nu}} (0 < \nu < 1)$$

An example: For the inverse power law $U(r)=r^{-(p-1)}$ $(p>2)\text{, } \gamma=\frac{p-5}{p-1}\text{, } \nu=\frac{1}{p-1}.$

Grad's angular cutoff assumption:

$$\int_0^{\pi/2} \sin\theta \tilde{b}(\cos\theta) \, d\theta < \infty.$$

- Landau collision (Landau, 1936):

$$\left(\frac{\partial f_s}{\partial t}\right)_c = \sum_{s'} Q(f_s, f_{s'}),$$

$$Q(f_1, f_2) = \frac{1}{m_1} \nabla_{\xi} \cdot \int_{\mathbb{R}^3} \Phi(\xi - \xi') \{ \frac{1}{m_1} f_1(\xi) \nabla_{\xi} f_2(\xi') - \frac{1}{m_2} f_2(\xi) \nabla_{\xi} f_1(\xi') \} d\xi',$$

$$\Phi(z) = |z|^{\gamma+2} (\mathbf{I} - \frac{z \otimes z}{|z|^2}) (\gamma \ge -3),$$

$$\gamma = -3 : \textbf{Coulomb potential}$$

Remark: Grazing limit: Boltzmann⇒Landau

3rd type (fluid plasma equations)

$$\begin{aligned} \partial_t n_s + \nabla \cdot (n_s v_s) &= 0, \\ m_s n_s (\partial_t v_s + v_s \cdot \nabla v_s) + \nabla P_s &= q_s n_s (E + \frac{v_s}{c} \times B) \\ &+ \sum_{s'} \nu_{ss'} \frac{m_s m_{s'} n_s n_{s'}}{m_s n_s + m_{s'} n_{s'}} (v_s - v_{s'}), \\ \frac{1}{c} \partial_t E - \nabla \times B &= -\frac{4\pi}{c} J, \end{aligned}$$

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$$\frac{1}{c}\partial_t E - \nabla \times B = -\frac{4\pi}{c}J,$$

$$\frac{1}{c}\partial_t B + \nabla \times E = 0,$$

$$\nabla \cdot E = 4\pi\rho, \quad \nabla \cdot B = 0,$$

$$\rho = \sum_s q_s n_s, \quad J = \sum_s q_s n_s v_s.$$

Euler-Maxwell system with/without relaxation

The Plasma Stability Problem

- Due to the collision AND particle-field interactive mechanism, a plasma usually relaxes to different kinds of profiles such as equilibrium states, periodic states, and wave patterns.
- Both physically and mathematically, it is an important task to understand the stability of those profiles.
- Stability theory addresses the following three questions:
 - Can the initial (small) perturbation of a given profile imply the global-in-time existence of solutions?
 - Will the solution converge to it? How fast for the rate of convergence?

If unstable, how to characterize the growth modes?

Remark: Problems without collisions are quite different (nonlinear effect and structure) !

- ▶ Molecule model: H. Weitzner (CPAM '12)
- ▶ Vlasov-Poisson system: Lemou-Mehats-Raphael (Inve. '11), Mouhot-Villani (Acta M. '11), ...

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► Euler-Maxwell system: Germain-Masmoudi (arXiv '11)

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Time-asymptotic stability of kinetic plasmas for general collision potentials

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Boltzmann's celebrated H-theorem

 $\partial_t f = Q(f, f) \Rightarrow$ (Physical) entropy increasing:

$$\frac{d}{dt}\int_{\mathbb{R}^3} d\xi \left\{-f\log f\right\} \ge 0.$$

This is a manifestation of the second law of thermodynamics.

Entropy takes the maximization at the Maxwellian

$$\mathbf{M} = \mathbf{M}_{[\rho, u, T]}(\xi) = \frac{\rho}{(2\pi T)^{3/2}} e^{-\frac{|\xi - u|^2}{2T}}.$$

 ρ : density, u: bulk velocity, T: temperature.

- ► L. Boltzmann himself predicted rapid convergence in large time to the Maxwellian due to the H-theorem. The "proof" was however held back by "analytical difficulties".
- ► Goal: prove convergence and convergence rate around the Maxwellian in the spatially non-homogeneous case.

Degeneration of H-theorem

$$\begin{aligned} \{\partial_t + \xi \cdot \nabla_x\}f &= Q(f, f) \Rightarrow \\ \frac{d}{dt} \int_{\Omega} dx \int_{\mathbb{R}^3} d\xi \left\{-f \log f\right\} \ge 0, \quad \Omega = \mathbb{R}^3 \text{ or } \mathbb{T}^3. \end{aligned}$$

H-theorem fails at the local Maxwellian

$$\mathbf{M}_{[\rho(t,x),u(t,x),T(t,x)]}(\xi).$$

In T³ case, a key tool to overcome the degeneration is the Poincare inequality:

$$\|\rho - \frac{1}{|\mathbb{T}^3|} \int_{\mathbb{T}^3} \rho \, dx \|_{L^2_x(\mathbb{T}^3)} \le C \|\nabla \rho\|_{L^2_x(\mathbb{T}^3)}$$

- In \mathbb{R}^3 case, the Poincare inequality fails.
- Idea: seek out the enough dissipative mechanisms for the components of the local Maxwellian

Degeneration of the electromagnetic field

For the Maxwell system in vacuum,

 $\partial_t E - \nabla \times B = 0, \quad \partial_t B + \nabla \times E = 0, \quad \nabla \cdot E = \nabla \cdot B = 0,$

the total energy is preserved at all time.

- Can the coupling with the kinetic equation imply a kind of the dissipative mechanism?
- Idea: again, seek out the enough dissipative mechanisms for the electromagnetic field with the understanding of the structure of the system

The Vlasov-Maxwell-Boltzmann/Landau system

$$\begin{split} f_{\pm} &= f_{\pm}(t, x, \xi) \geq 0 \text{ of two-species:} \\ \partial_t f_+ &+ \xi \cdot \nabla_x f_+ + (E + \xi \times B) \cdot \nabla_{\xi} f_+ = Q(f_+, f_+) + Q(f_+, f_-), \\ \partial_t f_- &+ \xi \cdot \nabla_x f_- - (E + \xi \times B) \cdot \nabla_{\xi} f_- = Q(f_-, f_+) + Q(f_-, f_-). \end{split}$$

It is coupled with the Maxwell system

$$\partial_t E - \nabla_x \times B = -\int_{\mathbb{R}^3} \xi(f_+ - f_-) d\xi,$$

$$\partial_t B + \nabla_x \times E = 0,$$

$$\nabla_x \cdot E = \int_{\mathbb{R}^3} (f_+ - f_-) d\xi, \quad \nabla_x \cdot B = 0.$$

The initial data in this system is given as

 $f_{\pm}(0, x, \xi) = f_{0,\pm}(x, \xi), \quad E(0, x) = E_0(x), \quad B(0, x) = B_0(x).$

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Previous results on VMB

Boltzmann collision term Q takes the hard sphere model:

$$B(\xi - \xi_*, \omega) = |(\xi - \xi_*) \cdot \omega|.$$

 $\blacktriangleright \ \Omega = \mathbb{T}^3$

- Global existence: Guo (IM, '03) (Energy method)
- Large-time behavior of solutions: Jang (ARMA, '09)
- $\blacktriangleright \ \Omega = \mathbb{R}^3$
 - Global existence: Strain (CMP, '06) (Use two-species' cancelation property to control *E* and pure time derivatives)
 - Large-time behavior of solutions: D.-Strain ('10) (Linearized analysis + bootstrap to the nonlinear equation)

! Unknown for non hard-sphere model !

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Previous results on VML

The only existing results concern the case of the absence of the variable magnetic field, i.e.

Vlasov-Poisson-Landau instead of VML

•
$$\Omega = \mathbb{T}^3$$
: Guo (JAMS, '12)
• $\Omega = \mathbb{R}^3$:

- D.-Yang-Zhao (arXiv '11): an application of the exponential weight
- Strain-Zhu (arXiv '12) and Yu (preprint '12): approach by Guo

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Wang (arXiv '12): pure energy method without linearized analysis

! Unknown in the case of VML !

Linearization (Carleman, Grad, ...)

- ► Define the perturbation u as $u = \mathbf{M}^{-1/2}(f \mathbf{M})$, $u = [u_+, u_-]$, $f = [f_+, f_-]$, $\mathbf{M} = \mathbf{M}_{[1,0,1]}(\xi)$.
- ▶ Boltzmann's H-theorem implies: $f_{\pm} \rightarrow \mathbf{M}$, $[E, B] \rightarrow 0$.
- The linearized system

$$\begin{cases} \partial_t u + \xi \cdot \nabla_x u - E \cdot \xi \mathbf{M}^{1/2} [1, -1] = \mathbf{L} u + g, \\ \partial_t E - \nabla_x \times B = -\langle [\xi, -\xi] \mathbf{M}^{1/2}, \{\mathbf{I} - \mathbf{P}\} u \rangle, \\ \partial_t B + \nabla_x \times E = 0, \\ \nabla_x \cdot E = \langle \mathbf{M}^{1/2}, u_+ - u_- \rangle, \quad \nabla_x \cdot B = 0, \\ [u, E, B]|_{t=0} = [u_0, E_0, B_0], \end{cases}$$

 $\ker \mathbf{L} = \operatorname{span} \left\{ [1,0]\mathbf{M}^{\frac{1}{2}}, [0,1]\mathbf{M}^{\frac{1}{2}}, [\xi_i, \xi_i]\mathbf{M}^{\frac{1}{2}} (1 \le i \le 3), [|\xi|^2, |\xi|^2]\mathbf{M}^{\frac{1}{2}} \right\}.$ $\models \text{ The local Maxwellian}$

$$\mathbf{P}_{\pm}u = \{a_{\pm}(t,x) + b(t,x) \cdot \xi + c(t,x)(|\xi|^2 - 3)\}\mathbf{M}^{\frac{1}{2}}.$$

Dissipation from L:

$$\int_{\mathbb{R}^3} u \cdot \mathbf{L} u \, d\xi \lesssim - \int_{\mathbb{R}^3} \nu(\xi) |\{\mathbf{I} - \mathbf{P}\} u|^2 \, d\xi$$

- ► Collision frequency: $\nu(\xi) = \langle \xi \rangle^{\gamma}$; **P**: projection from L_{ξ}^2 to ker **L**
- A summary of possible difficulties:
 - The dissipation of Pu is missing: The local Maxwellian is dispersive in the whole space due to the degeneration of L ! (Hypocoercivity: Villani)
 - If γ < 1 then how to control a nonlinear term which grows in large |ξ| at least linearly?
 - If γ < 0 then how to control 1st-order velocity derivative of the linear transport term ξ · ∇_xu?
 - If γ < 0 is much smaller then how to control the nonlinear transport term E · ∇_ξu provided that the velocity differentiation needs the extra velocity weight?

Dissipation of $\mathbf{P}u$

Observation (Grad, Kawashima, Liu-Yu, Guo, D.-Strain '10):

Find dissipation from the dynamics of the local Maxwellian?

$$\begin{aligned} \partial_t a_{\pm} + \nabla_x \cdot b + \nabla_x \cdot \langle \xi \mathbf{M}^{1/2}, \{ \mathbf{I}_{\pm} - \mathbf{P}_{\pm} \} u \rangle &= 0, \\ \partial_t [b_i + \langle \xi_i \mathbf{M}^{1/2}, \{ \mathbf{I}_{\pm} - \mathbf{P}_{\pm} \} u \rangle] + \partial_i (a_{\pm} + 2c) \mp E_i \\ + \nabla_x \cdot \langle \xi \xi_i \mathbf{M}^{1/2}, \{ \mathbf{I}_{\pm} - \mathbf{P}_{\pm} \} u \rangle &= 0, \\ \partial_t \left[c + \frac{1}{6} \langle (|\xi|^2 - 3) \mathbf{M}^{1/2}, \{ \mathbf{I}_{\pm} - \mathbf{P}_{\pm} \} u \rangle \right] + \frac{1}{3} \nabla_x \cdot b \\ &+ \frac{1}{6} \nabla_x \cdot \langle (|\xi|^2 - 3) \xi \mathbf{M}^{1/2}, \{ \mathbf{I}_{\pm} - \mathbf{P}_{\pm} \} u \rangle = 0, \end{aligned}$$

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These equations are NOT closed!

To close the system of the local Maxwellian, we also need to study the high-order moment equations:

$$\begin{aligned} \partial_t [\Theta_{ii}(\{\mathbf{I}_{\pm} - \mathbf{P}_{\pm}\}u) + 2c] + 2\partial_i b_i &= \Theta_{ii}(l_{\pm}), \\ \partial_t \Theta_{ij}(\{\mathbf{I}_{\pm} - \mathbf{P}_{\pm}\}u) + \partial_j b_i + \partial_i b_j + \nabla_x \cdot \langle \xi \mathbf{M}^{1/2}, \{\mathbf{I}_{\pm} - \mathbf{P}_{\pm}\}u \rangle \\ &= \Theta_{ij}(l_{\pm}), \quad i \neq j, \\ \partial_t \Lambda_i(\{\mathbf{I}_{\pm} - \mathbf{P}_{\pm}\}u) + \partial_i c = \Lambda_i(l_{\pm}). \end{aligned}$$

Here, the high-order moment functions are defined by

$$\Theta_{ij}(u_{\pm}) = \langle (\xi_i \xi_j - 1) \mathbf{M}^{1/2}, u_{\pm} \rangle, \Lambda_i(u_{\pm}) = \frac{1}{10} \langle (|\xi|^2 - 5) \xi_i \mathbf{M}^{1/2}, u_{\pm} \rangle,$$

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and L_{\pm} are defined in terms of $\{I - P\}u$.

 \blacktriangleright For $\gamma \geq 0,$ there is a time-frequency functional $\mathcal{E}(t,k)$ such that

$$\mathcal{E}(t,k) \sim \|\hat{u}\|_{L_{\xi}^{2}}^{2} + |\hat{E}|^{2} + |\hat{B}|^{2},$$

and

$$\partial_t \mathcal{E}(t,k) + \frac{\lambda |k|^2}{(1+|k|^2)^2} \mathcal{E}(t,k) \le 0, \quad \forall t \ge 0, \ k \in \mathbb{R}^3.$$

Remark:

- ► The inequality seems terrible to prove decay rates because the "dissipative term" goes to zero as |k| → ∞.
- It is an essential "regularity-loss" feature for the VMB system, not a deficiency of our approach; see D. (Eigenvalue analysis of damped Euler-Maxwell, '11), Hosono-Kawashima (M3AS '06), Ueda-D.-Kawashima ('11):

$$\lambda(ik) \sim -\frac{1}{|k|^2} \pm i|k| \quad (|k| \to \infty).$$

Case when $\gamma < 0$: D. (arXiv '12)

The situation becomes more subtle for $\gamma < 0$; see Strain ('11) and D.-Yang-Zhao ('11)

• Let
$$w = w(\xi) = \langle \xi \rangle^{\frac{\gamma+2}{2}}$$
 for Landau.

Derive

$$\partial_t M_\ell(t,k) + \kappa D_\ell(t,k) \le 0,$$

with

$$\begin{split} M_\ell(t,k) &= \left\| \hat{u} \right\|_{L^2}^2 + \left| [\hat{E}, \hat{B}] \right|^2 + \kappa_0 \, \Re \, \mathcal{E}^{\text{int}}(t,k) \\ &+ \kappa_2 \left| w^\ell \{ \mathbf{I} - \mathbf{P} \} \hat{u} \right|_{L^2} \, \chi_{|k| \leq 1} + \frac{\kappa_1}{1 + |k|^2} \left| w^\ell \hat{u} \right|_{L^2}^2 \, \chi_{|k| \geq 1}, \\ D_\ell(t,k) &= \left| \{ \mathbf{I} - \mathbf{P} \} \hat{u} \right|_{\mathbf{D}}^2 + \frac{1}{1 + |k|^2} \left| w^\ell \{ \mathbf{I} - \mathbf{P} \} \hat{u} \right|_{\mathbf{D}}^2 \\ &+ \frac{|k|^2}{1 + |k|^2} (|\widehat{a_+ + a_-}|^2 + |\hat{b}|^2 + |c|^2) + |\widehat{a_+ - a_-}|^2 \\ &+ \frac{1}{1 + |k|^2} |\hat{E}|^2 + \frac{|k|^2}{(1 + |k|^2)^2} |\hat{B}|^2, \quad \rho(k) = \frac{|k|^2}{(1 + |k|^2)^2}. \end{split}$$

Make the time weighted estimate

$$M_{\ell}(t,k) \lesssim [1 + \epsilon \rho(k)t]^{-J} M_{\ell+J+p-1}(0,k)$$

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Nonlinear perturbation theory for general collisional potentials

- Mathematically, when there is an external force, it is highly nontrivial to generalize existing results to the case of non hard-sphere model, which is also of physical importance!
- A progress was made by Guo (JAMS, '11):
 - \blacktriangleright the Landau collision with the Coulomb potential $(\gamma=-3)$
 - the potential force $E = -\nabla \phi$ with vanishing B = 0:

$$(\xi \cdot \nabla_x u + \nabla_x \phi \cdot \xi u) e^{\phi} = \xi \cdot \nabla_x (e^{\phi} u).$$

 $\blacktriangleright \ \Omega = \mathbb{T}^3$

It is difficult to deal with the non-potential force !!!

- A completely different approach was developed by D.-Yang-Zhao (arXiv '11)
 - A new dissipative mechanism due to the introduction of the time-velocity dependent weight

 $\exp\{\lambda\langle\xi\rangle^q/(1+t)^\theta\}$

$$\Rightarrow \partial_t e^{\frac{\lambda\langle\xi\rangle^q}{(1+t)^{\theta}}} = -\lambda\theta \frac{\langle\xi\rangle^q}{(1+t)^{1+\theta}} e^{\frac{\lambda\langle\xi\rangle^q}{(1+t)^{\theta}}}.$$

- The approach that we developed can apply to
 - Landau or Boltzmann

•
$$\Omega = \mathbb{R}^3$$
 or \mathbb{T}^3

- For the Boltzmann with most of values of γ : angular cutoff or non-cutoff
- Maxwell system (non-potential force) can be included!

Main results:

Global classical solutions near a global Maxwellian uniquely exist and time asymptotically tend to the Maxwellian with some rates for the cases of

- ▶ D.-Yang-Zhao ('11): Vlasov-Poisson-Boltzmann, angular cutoff with $-2 \le \gamma \le 1$
- ▶ D.-Liu ('11): Vlasov-Poisson-Boltzmann, angular non cutoff with $-3 < \gamma < -2$ and $1/2 \le s < 1$

▶ D. (arXiv '12): Vlasov-Maxwell-Landau, soft potentials $-3 \le \gamma < -2$ including the Coulomb $\gamma = -3$

Idea in the proofs:

► Find an energy functional $\mathcal{E}(t)$ and its time-weighted norm X(t) such that

 $X(t) \lesssim Y_0 + [X(t)]^2.$

- ▶ To control the term $E \cdot \nabla_{\xi} u$ and $E \cdot \xi u$, the time-decay of E is needed. Thus, Y_0 generally includes L^1 -norm of initial data. Note that Y_0 needs to be small enough to ensure the global-in-time bound by the continuity argument.
- ▶ To balance an estimate on both $E \cdot \nabla_{\xi} u$ and $\xi \cdot \nabla_{x} u$, γ can NOT be too small in the cutoff case. However, in the non cutoff case, since

$$\int u \mathbf{L} u \, d\xi \lesssim -\int \langle \xi \rangle^{\gamma+2s} |\{\mathbf{I} - \mathbf{P}\} u|^2 \, d\xi - \{\cdots\},$$

we may require that $\gamma + 2s$ need not be too small.

Idea in the proofs (cont.):

- ► To deal with the degeneration of $\nu(\xi)$ for soft potentials, choose $\mathcal{E}(t)$ in the way that
 - ▶ higher the differentiation order is, the order of velocity weights is lower, for instance, consider (∂^β_α = ∂^β_x∂^α_ξ, |α| + |β| = N)

$$\iint \partial_{\alpha}^{\beta} Q(u, u) \cdot w_{\alpha, \beta, \ell}^{2}(t, \xi) \partial_{\alpha}^{\beta} u \, dx d\xi.$$

- ► To deal with the degeneration of the Maxwell equations, choose *X*(*t*) in the way that
 - ▶ higher the order of $\mathcal{E}_N(t)$ is, the rate of its time weights is lower;
 - ▶ the highest-order energy norm $\mathcal{E}_N(t)$ may increase in time! For instance, consider

$$\iint \partial_{\alpha}^{\beta}[(B \times \xi) \cdot \nabla_{\xi} u] \cdot w_{\alpha,\beta,\ell}^{2}(t,\xi) \partial_{\alpha}^{\beta} u \, dx d\xi.$$

Idea in the proofs (cont.):

A trouble occurs to the estimate on

 $\iint \partial_{\beta}^{\alpha} (E \cdot \xi \mathbf{M}^{1/2}) \cdot w_{\alpha,\beta,\ell}^2(t,\xi) \partial_{\beta}^{\alpha} u \, dx d\xi \quad (|\alpha| + |\beta| = N)$

No dissipation for $\int |\nabla_x^N E|^2$! Use an idea from Hosono-Kawashima (M3AS '06) to deduce

$$\frac{d}{dt} [(1+t)^{-\epsilon_0} \mathcal{E}_{N_1}(t)] + \kappa (1+t)^{-\epsilon_0} \mathcal{D}_{N_1}(t)
+ \frac{\epsilon_0}{(1+t)^{1+\epsilon_0}} \mathcal{E}_{N_1}(t) \lesssim (1+t)^{-\epsilon_0} \times \{h.o.t.\}.$$

from the basic energy inequality without any weight

$$\frac{d}{dt}\mathcal{E}_{N_1}(t) + \kappa \mathcal{D}_{N_1}(t) \lesssim h.o.t..$$

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Problems for the future:

- Bounded domain
- Eigenvalue analysis and the spectrum

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Thank you !

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